Imaging via photon-photon correlation of X-ray fluorescence



Fabian Trost SPB/SFX

Hamburg, 01/24/2024





Introduction



Fabian Trost, 01/24/2024



X-ray Flouresence

- Fluorescence is incoherent.
 - It averages to a flat intensity distribution,
 Where all structural information is lost.
- When measured within timescales of the coherence time a speckle pattern is obtained.



(speckles origin from interference)



average over 1 000 000 exposures of Cu K_{α} -fluorescence (no structure, just flat intensity distribution)

Photon-photon correlation

Think of a triple slit:



Photon-photon correlation

European XFEL

Think of a triple slit, with a random phase shift after each exposure:



$$g^{(2)} = \frac{\langle I(k)I(k+q)\rangle_k}{\langle I(k)\rangle_k^2}$$

Photon-photon correlation

- Fluorescence is incoherent.
 - Averages to flat intensity distribution.
- When measured within timescales of the coherence time τ_c , a speckle pattern is obtained.
- The speckle pattern encodes the "coherent pattern", which can be obtained

via autocorrelation of the measured photons:
$$g^{(2)}(\vec{q}) = \frac{\langle I(\vec{k}) I(\vec{k}+\vec{q}) \rangle_{\vec{k}}}{\langle I(\vec{k}) \rangle_{\vec{k}}^2}$$
.

$$g^{(2)}(\vec{q}) = 1 + \beta \left| g^{(1)}(\vec{q}) \right|^2 \text{ encodes the emitter distribution } \rho: g^{(1)}(\vec{q}) = \frac{\mathfrak{F}[\rho(\vec{r})](\vec{q})}{\mathfrak{F}[\rho(\vec{r})](0)}$$



averaged fluorescence

0.2

Photon-photon correlation (short history)

Robert Hanbury Brown and Richard Twiss used intensity interferometry to measure the diameter of stars.

Intensity correlation signal must origin from multiphoton interference.

Contradicted Paul Diracs view that "each photon [...] only interferes with itself"

Explanation by Ugo Fano: photon paths are not distinguishable and therefore interfere.

> [U. Fano, "Quantum theory of interference effects in the mixing of light from phaseindependent sources." American Journal of Physics 29.8 (1961): 539-545.]

[Sky & Telescope, vol. 28, pp. 2-7, 1964]



10

Base-line, d (metres)

12

14

[R. Hanbury Brown, R. Twiss, Nature, Vol 178 Springer p. 1046-1048 (1956)] Comparison between the values of the normalized correlation coefficient $\Gamma^2(d)$ observed from Sirius and the theoretical values for a star of angular diameter 0.0063".

[Paul Dirac, The principles of

quantum mechanics, 1930]



Photon-photon correlation imaging (motivation)

- Element sensitivity.
- High cross-section.
- Fluorescence is isotropic.
- Translation invariant.
- 3D information and higher resolution.
- Can be used for FEL pulse length determination.





Photon-photon correlation imaging (motivation)



usable signal

Usable signal - visibility factor β

Not all fluorescence photons can interfere. \rightarrow independent coherent modes *M*.

 $g^{(2)}(\vec{q})$

 \rightarrow visibility factor $\beta = 1/M$ modulates the "usable signal".

Fluorescence is unpolarized

We can not distinguish $K_{\alpha,1}$ and $K_{\alpha,2}$

Finite excitation pulse length

$$\rightarrow \beta_{\text{lines}} \approx \frac{5}{9}$$
.

 $\rightarrow \beta_{\text{pol}} = \frac{1}{2}$.

offset

$$\rightarrow \beta_{\text{pulse}} \approx \frac{\tau_c}{1.5 T}$$

→ Expected visibility
$$\beta \approx \frac{5 \tau_c}{27 T}$$
.





Signal-to-noise ratio - photon statistics

Consider many (classical) emitters with random phases. Theirs intensity follow an exponential distribution

 $P_{\rm Exp}(I|\mu) = \frac{1}{\mu} e^{-\frac{I}{\mu}}.$



We measure discrete photons (Poisson distributed) \rightarrow Bose-Einstein distribution $P_{\text{BE}}(n|\mu) = \frac{1}{1+\mu} + \left(\frac{\mu}{1+\mu}\right)^n$.

Independent modes $M = \beta^{-1}$ yields a negative binomial distribution $P_{\text{NB}}(n|\mu, M) = \frac{M^M \mu^n \Gamma(M+n)}{(M+\mu)^{M+n} n! \Gamma(M)}$, with $\mathbb{E}(n) = \mu = \langle I \rangle$ and $\text{Var}(n) = \mu + \beta \mu^2$.

Signal-to-noise ratio - photon statistics

Photon statistics: negative binomial

$$\blacksquare \mathbb{E}(n) = \mu = \langle I \rangle, \operatorname{Var}(n) = \mu + \beta \mu^2.$$

Assuming uncorrelated intensities:

SNR =
$$\frac{\beta |g^{(1)}(\vec{q})|^2 \mu^2 \sqrt{N_p C(\vec{q})}}{\sqrt{(\beta^2 + 4\beta)\mu^4 + (4 + 2\beta)\mu^3 + \mu^2}}$$



$(\vec{q}) \Big|^2$. $(\vec{s}) \Big|^2$ $(\vec{q}) \Big|^2$ $(\vec{q}) \Big|^2$ $(\vec{s}) \Big|^{150} - 200 - 2$

0





Signal-to-noise ratio - complexity

The "usable signal" is on top of an offset $g^{(2)}(\vec{q}) = 1 + \beta |g^{(1)}(\vec{q})|^2$.

However, the offset is noisy.

Remember $\sigma_{AC} = \sqrt{(\beta^2 + 4\beta)\mu^4 + (4 + 2\beta)\mu^3 + \mu^2}$.

- For example crystals:
 - Bragg peaks are on top of the offset.
 - Offset is proportional to $N_{\rm uc}^2$.
 - "Usable signal" is proportional to N_{uc}.
 ▶ Signal-to-offset ratio is ∝ 1/N_{uc}.

Signal-to-noise ratio - complexity

Crystals: more unit cells lead to a poorer SNR, always!



The offset grows usually faster than the usable signal with more resolvable emitters.

 $|\mathfrak{F}[\rho(\vec{r})](0)| \ge |\mathfrak{F}[\rho(\vec{r})](\vec{q})| \text{ if } \rho(\vec{r}) \ge 0 \ \forall \vec{r} \Rightarrow \ \left|g^{(1)}(\vec{q})\right|^2 \le 1.$

Further reading: Trost et al. 2020 New J. Phys. 22 083070





Experimental demonstration – Cu-target

EuXFEL provided trains with 444ns spaced pulses, each powerful enough to drill through the Cu foil.

- The Cu target is replaced after each pulse, by spinning the disk at around 4500 rpm.
- We were able to take a total of approximately 600 million exposures during the beamtime (3.1PB raw data)





Experimental demonstration – imaging via photon-photon correlation of X-ray fluorescence

We measured 58 million patterns with the phase-grating.

Averaged photon counts shows a flat intensity distribution. $\langle I \rangle = 0.0077$.

Pattern-wise autocorrelation reveals fringes.





Experimental demonstration – imaging via photon-photon correlation of X-ray fluorescence

Fabian Trost, 01/24/2024

Phasing reveals the emitter distribution.



Integrated usable signal was around 33% with the grating in comparison to the case without grating.





$g^{(2)} - 1$ without phase grating

FEL pulse and focus characterization

Photon-photon correlation can be used for focus finding and characterization.



Retrieved focus was (640 ± 40) nm × (480 ± 30) nm FWHM.

Each data point consist of between 750 000 and 3 000 000 patterns, acquired within 5 - 20 min.

▶ Could be reduced to ~10 000 patterns for 10 times more photons \rightarrow less than 5 sec.

Insensitive to beam jitter.

FEL pulse and focus characterization

Signal fitting also yields the visibility factor.





We have determined $\beta = (0.018 \pm 0.002)$, with $\beta \approx 0.185 \frac{0.6\text{fs}}{T}$:

▶ retrieved pulse duration $T = (6.2 \pm 0.8)$ fs (FWHM).

Further reading: *Trost et al. Phys. Rev. Lett.* 130, 173201



[also demonstrated by Inoue et al., see J. Synchrotron Rad. 26, 2050–205 (2019).]

Imaging via photon-photon correlation of X-ray fluorescence

Summary and Outlook

- Imaging via correlation of X-ray fluorescence is possible.
 can be used for robust focus measurements.
- Imaging of complex emitter distributions is complicated by a loss in SNR.
- Visibility is dependent on the excitation pulse duration.- can be used to determine the pulse length.

Next steps:

- Demonstrate single particle imaging via photon-photon correlation.
- Implement photon-photon correlation imaging as a robust tool for beam characterization.
- Advent of attosecond pulses on FELs will open new possibilities for X-ray fluorescence correlation imaging.







Fabian Trost, 01/24/2024

Thank you!

Henry Chapman, Holger Fleckenstein, Sasa Bajt, Mauro Prasciolu, Oleksandr Yefanov, Tjark Delmas, Martin Domaracky, Jerome Carnis, Chufeng Li, Miriam Barthelmess, Lukas Dresselhaus (CFEL DESY)

Kartik Ayyer, Tamme Wollweber, Abhishek Mall, Zhou Shen, Yulong Zhuang (MSPD)

Joachim von Zanthier, Stefan Richter, Sebastian Karl (Univ. Erlangen Nürnberg, Germany)

Ralf Röhlsberger (DESY and Jena)

Nic Timneanu, Carl Caleman, Sebastian Cardoch, Kajwal Kumar (U. Uppsala)

Anders Madsen, Johannes Möller, Alexey Zozulya, Roman Shayduk, Wei Lu, Felix Brauße, Bertram Friedrich, Ulrike Boesenberg, Ilia Petrov, Sergey Tomin, Marc Guetg (EuXFEL)

contact Fabian Trost fabian.trost@xfel.eu +49 040 8998 6780