

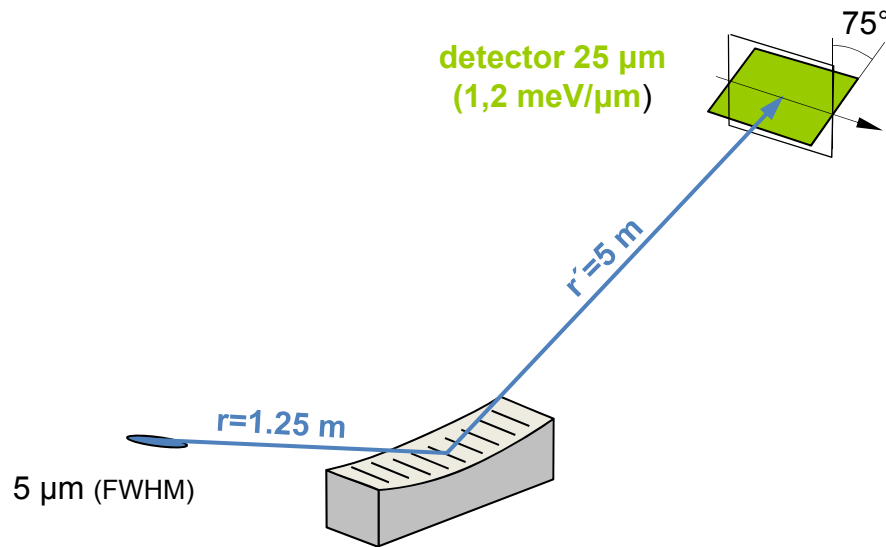


Rolf Follath

RIXS spectrometer designs

- **V**LS Spectrometer
- **P**lane Grating Spectrometer

VLS-Spectrometer



$$R_G = 108\text{ m}, L=150\text{ mm}$$

$$N = 3200\text{ l/mm}$$

$$v_1 = 2.29\text{e-}4\text{ 1/mm,}$$

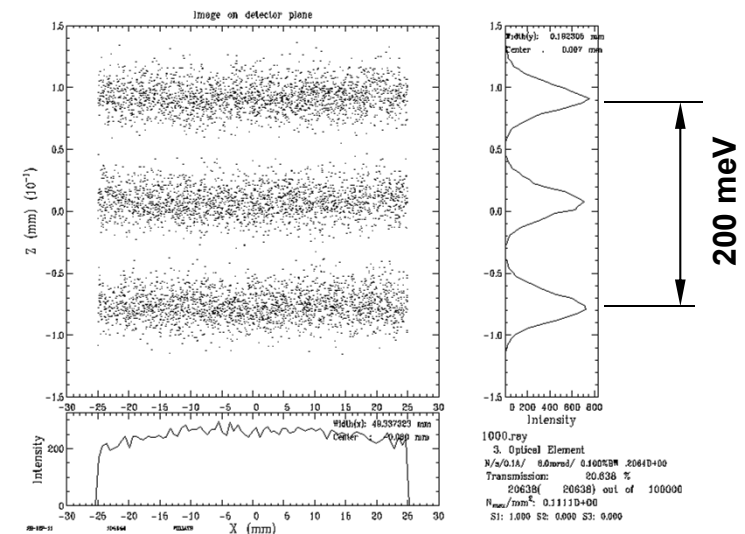
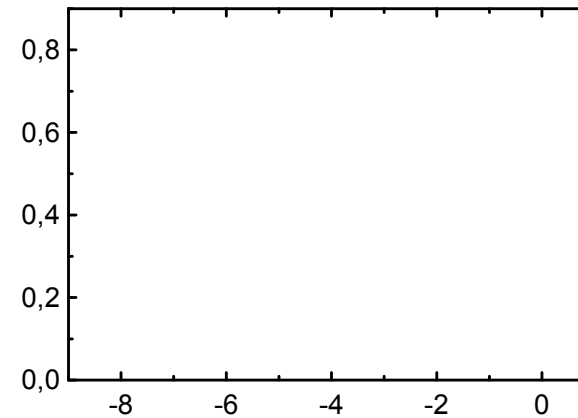
$$v_2 = 2.07\text{e-}8\text{ 1/mm}^2$$

$$\sigma = 0.25\text{ }\mu\text{rad}$$

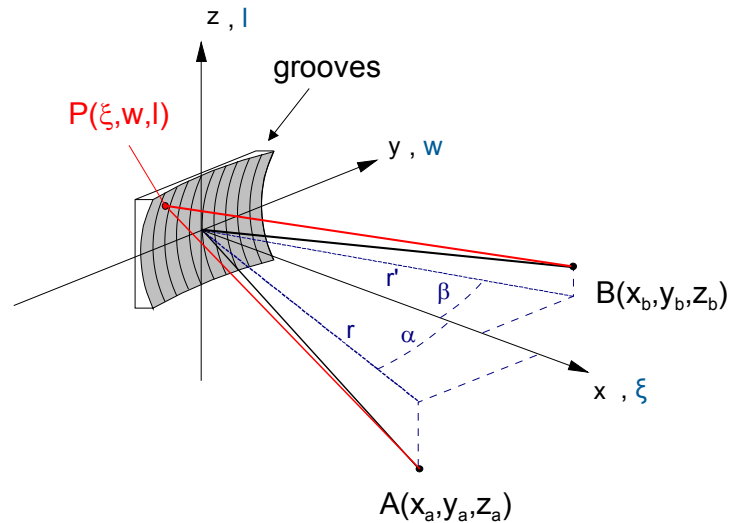
Acceptance : detector x 2 mrad²

Calculations for detector width of 50 mm

Resolution \approx 30 meV @ 1000 eV



Path function



R : Radius of curvature
 r : distance source to grating
 r' : distance grating to detector
 α : incident angle
 β : diffraction angle

Optical Path function

$$F = \overline{AP} + \overline{PB} + N(w) \cdot \lambda$$

$N(w)$: Number of lines between 0 and w

Taylor expansion of F in the coordinates of the grating surface w, l

$$F = F_{00} + F_{10} w + \frac{1}{2} F_{20} w^2 + \frac{1}{2} F_{02} l^2 + \frac{1}{2} F_{30} w^3 + \frac{1}{2} F_{12} w l^2 + \dots$$

VLS operation parameter

For spherical gratings with variable line separation $d(w) = d_0 (1 + \nu_1 \cdot w + \nu_2 \cdot w^2 + \nu_3 \cdot w^3 + \dots)$

$$F_{10} = m \frac{\lambda}{d_0} - (\sin \alpha + \sin \beta) \quad \text{grating equation}$$

$$F_{20} = \left(\frac{\cos^2 \alpha}{r} - \frac{\cos \alpha}{R} \right) + \left(\frac{\cos^2 \beta}{r'} - \frac{\cos \beta}{R} \right) - \nu_1 m \frac{\lambda}{d_0} \quad \text{meridional focus (xy-plane)}$$

$$F_{30} = \left(\frac{\cos^2 \alpha}{r} - \frac{\cos \alpha}{R} \right) \frac{\sin \alpha}{r} + \left(\frac{\cos^2 \beta}{r'} - \frac{\cos \beta}{R} \right) \frac{\sin \beta}{r'} + \frac{2}{3} (\nu_1^2 - \nu_2) m \frac{\lambda}{d_0} \quad \text{coma}$$

Three conditions to fulfill,

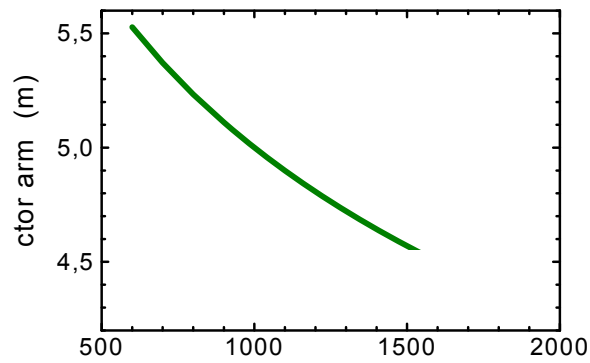
Use two angles α, β and r' , the distance between grating and detector

Detector angle

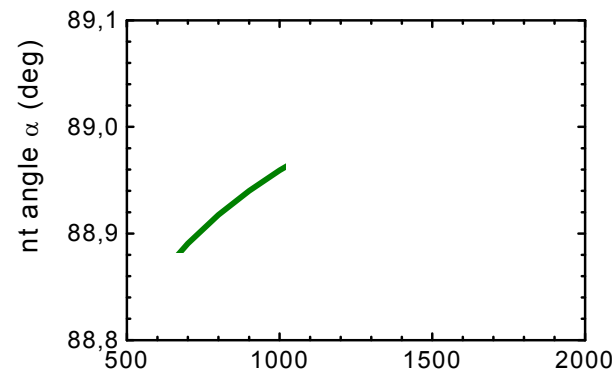
$$\tan \phi = \frac{\cos \beta}{r' \left(\frac{\tan \beta}{R} - \nu_1 \right) - 2 \sin \beta}$$

VLS system settings for α, β, r'

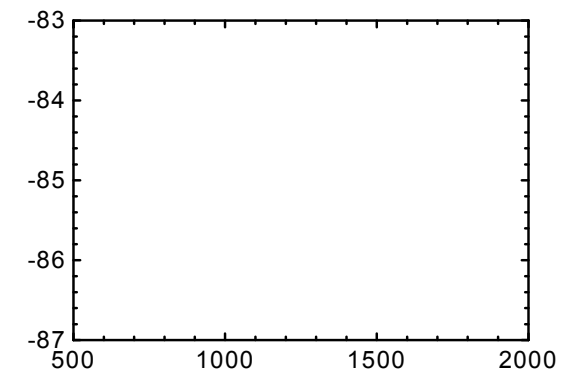
detector arm



entrance angle α



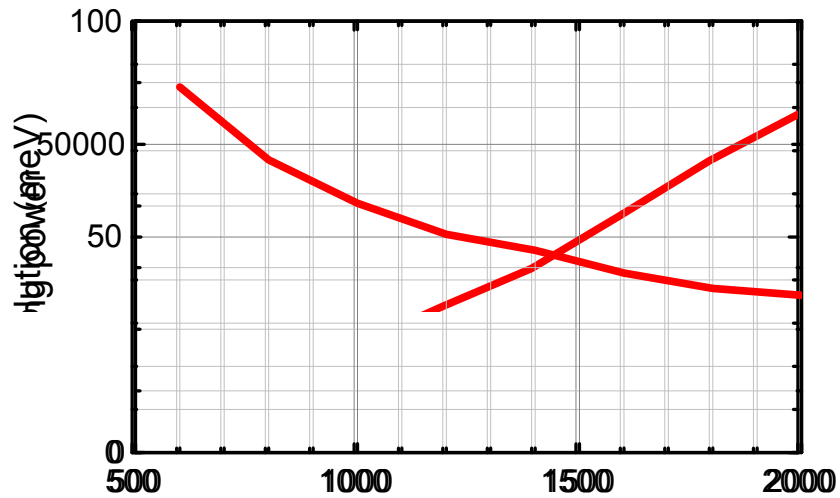
diffraction angle



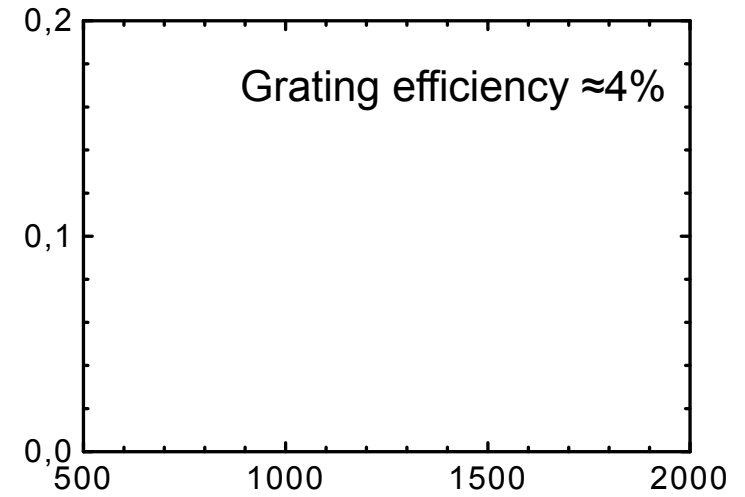
Detector position variation : hor: **1 m**, vert. **0,8 m**

VLS performance

Energy resolution



Effective vert.. acceptance



Hor. acceptance determined by detector

Add horizontally collecting mirror or use grating to focus sagittally

Contributions to energy resolution

Energy	1000 eV	(600+10) eV
Abberations	1,8 meV ✓	4.4 meV ✓
Slope errors	17 meV	8.3 meV
Source size	19 meV	8.4 meV
Total	26 meV	11 meV

Limited by source and slope errors

Alignment:

Detector position +/- 2 mm	30 meV	15 meV
+/- 4 mm	46 meV	24 meV
Source distance +/- 2 mm	26 meV	12 meV
+/- 4 mm	34 meV	16 meV

Alignment on the mm-scale

Contributions to spot size on detector

Energy	1000 eV	(600+10) eV
Abberations	0.4 μm ✓	6 μm ✓
Slope errors	3.7 μm	11.2 μm
Source size	4.2 μm	11.4 μm
Total	5.6 μm	15 μm

Limited by source and slope errors

Alignment:

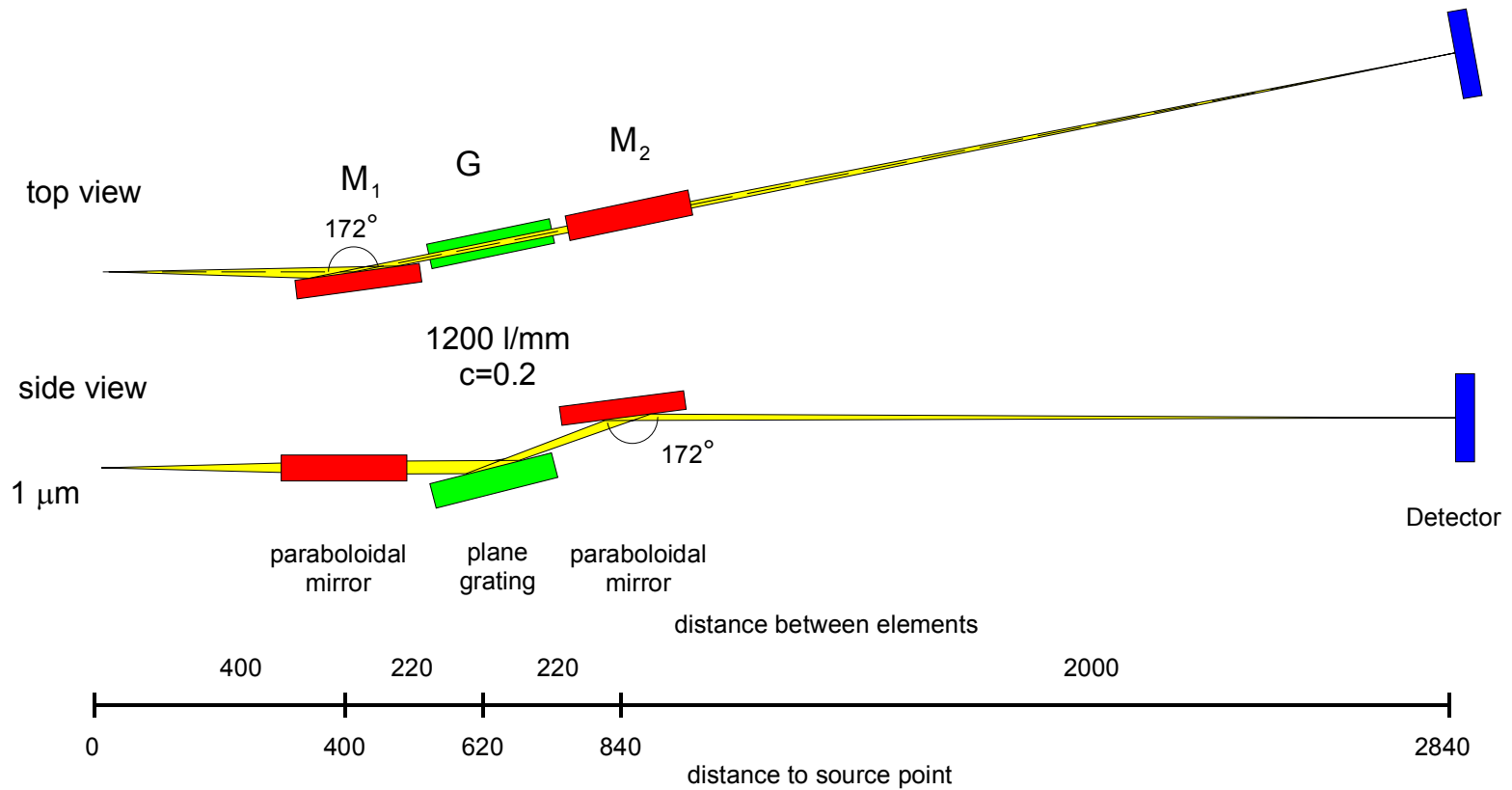
Detector position +/- 2 mm	6.4 μm	20 μm
+/- 4 mm	10 μm	32 μm
Source distance +/- 2 mm	5.7 μm	16 μm
+/- 4 mm	7.4 μm	22 μm

Alignment on the mm-scale

Plane Grating Spectrometer for UE49-SGM at BESSY II

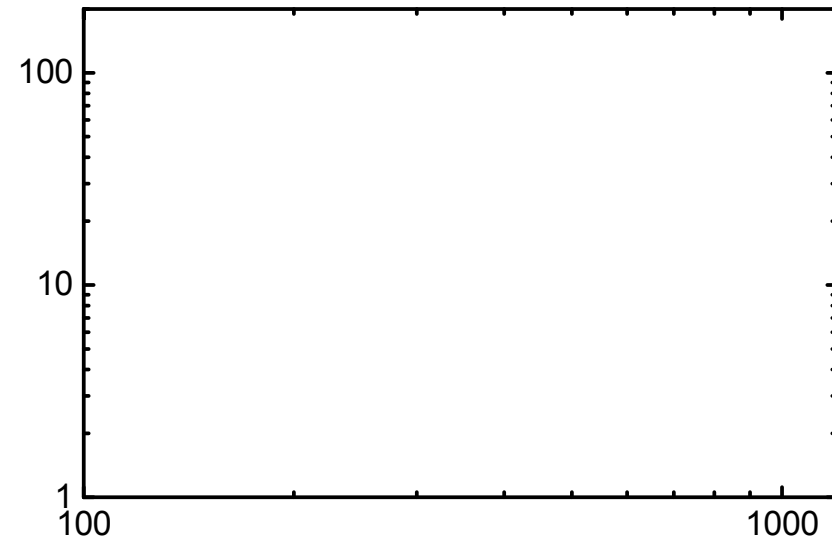
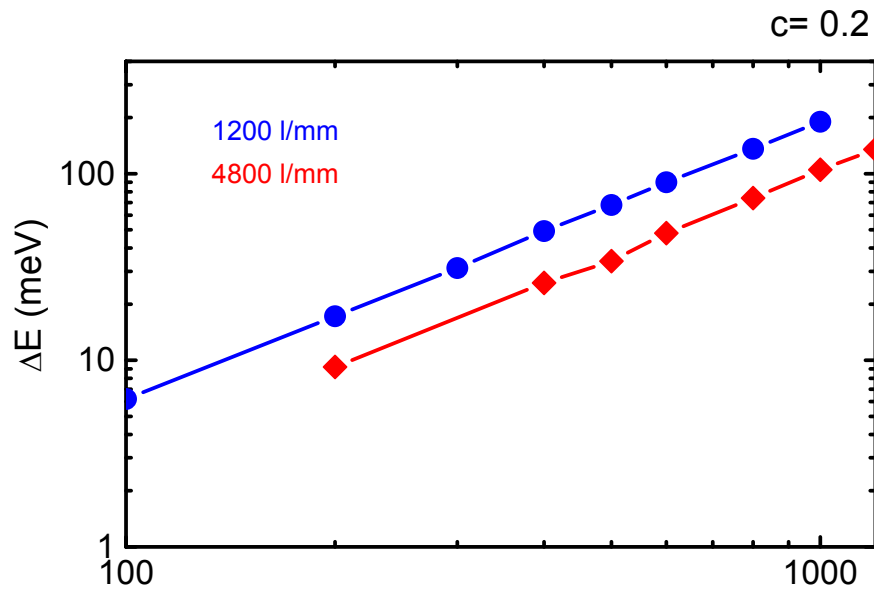
Following an instrument design by J.Nordgren group*

Modified design for BESSY II



* M. Agaker: Nucl. Instrum. and Methods, A **601** 213 (2009)

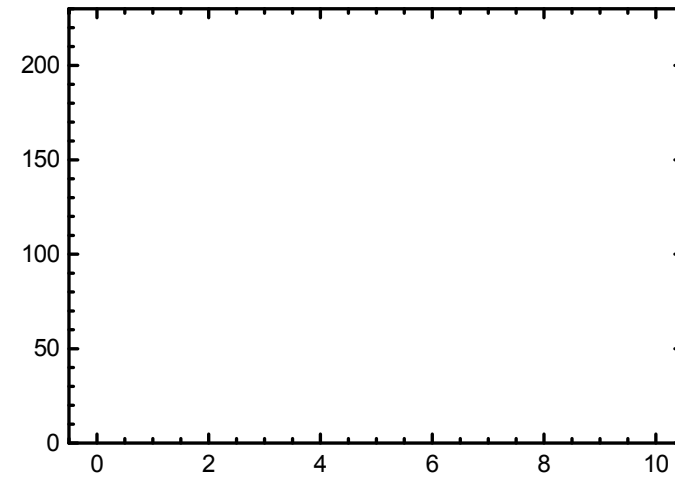
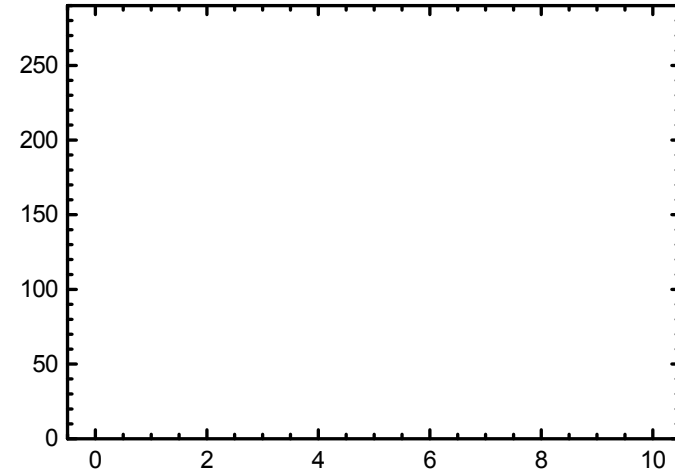
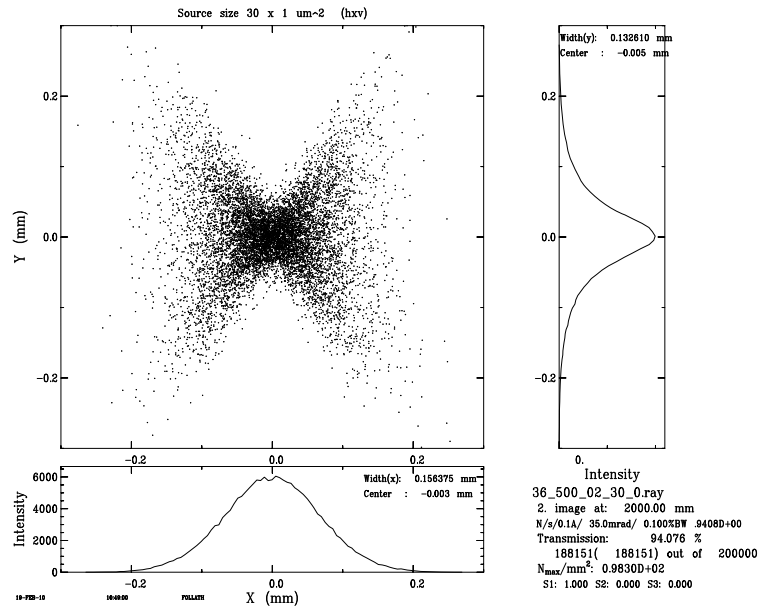
Performance



- Effective transmission = Transmission • Geometrical Acceptance
 - Transmission includes reflectance and grating efficiency
 - Geometrical acceptance: 35 mrad x 40 mrad
- Slope errors included
- Source size 1 μm (achieved in Oct. 2011)

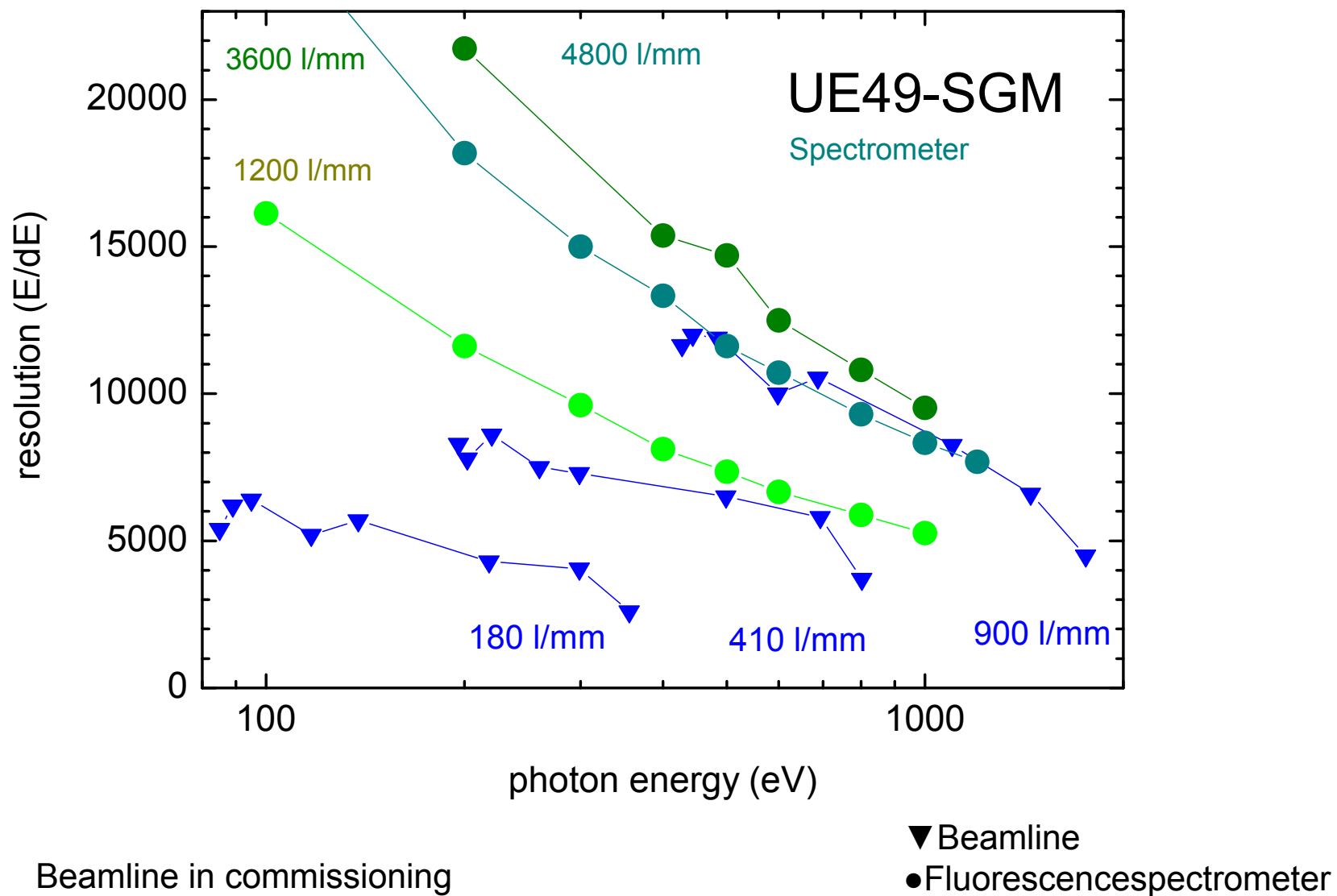
Source size effect

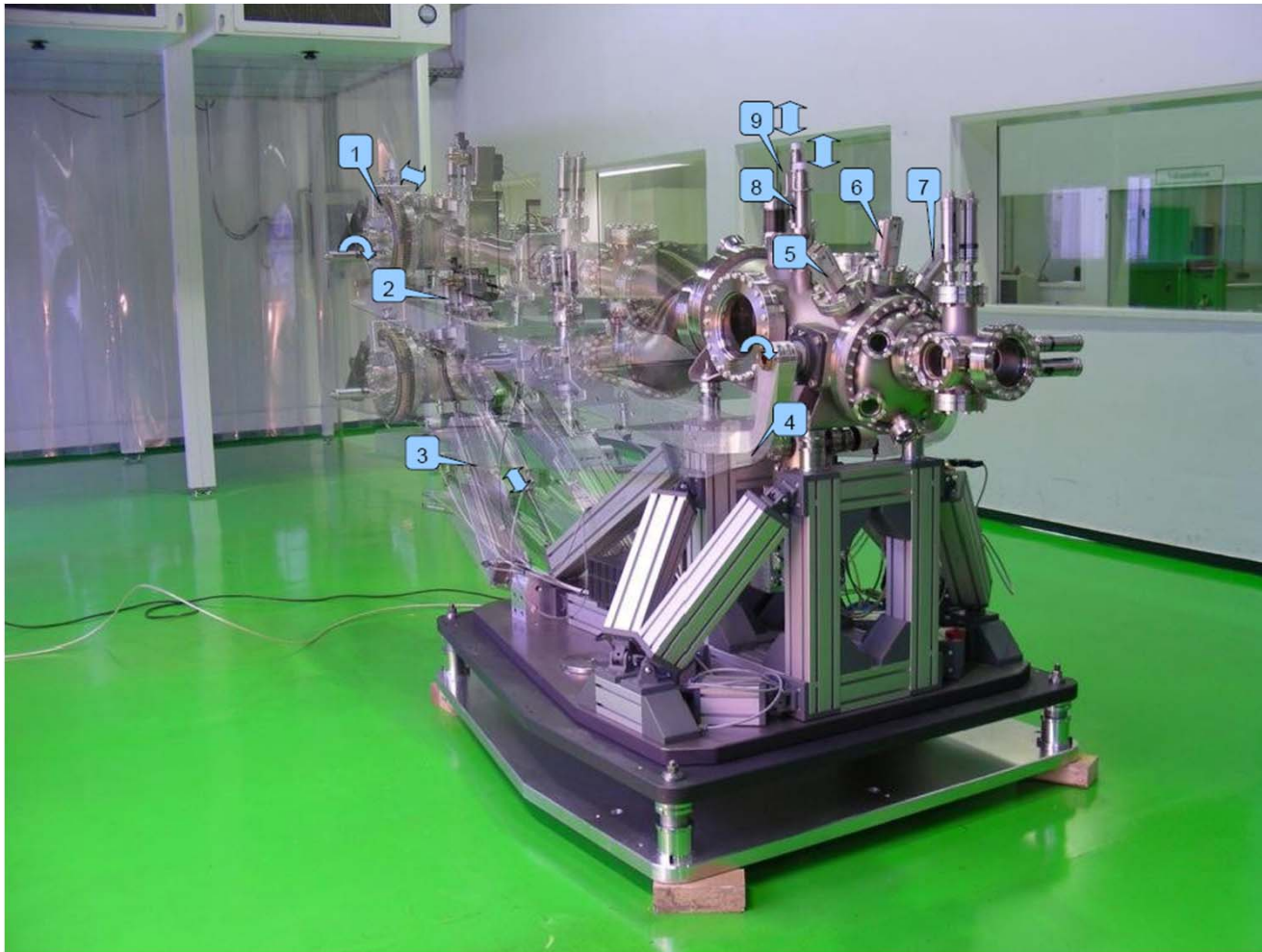
Image for source size of 30 x 1 μm



Center part has better resolution
=> apertures?

Resolving power Beamline vs. spectrometer







Thank You